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# **Reconstructing the total shortening history of the NW Himalaya**

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[1] The onset of India-Asia contact can be dated with both biostratigraphic analysis of syn-collisional sedimentary successions deposited on each side of the Indus Suture zone, and by

radiometric dating of Indian crustal rocks which have undergone subduction to great depths in the earliest subduction-collision stages. These data, together with paleomagnetic data show that the initial contact of the Indian and Asian continental margins occurred at the Paleocene/Eocene boundary, corresponding to  $55 \pm 2$  Ma. Such dating, which is consistent with all available geological evidence, including the record of magnetic anomalies in the Indian ocean and decrease of magmatic activity related to oceanic subduction can thus be considered as accurate and robust. The sedimentary record of the Tethys Himalaya rules out obduction of oceanic allochthons directly onto the Indian continental margin during the Late Cretaceous. The commonly inferred Late Cretaceous ophiolite obduction events may have thus occurred in intra-oceanic setting close to the Asian margin before its final emplacement onto the India margin during the Eocene. Granitoid and sedimentary rocks of the Indian crust, deformed during Permo-Carboniferous rifting, reached a depth of some 100 km about 1 Myr after the final closure of the Neo-Tethys, and began to be exhumed between 50 and 45 Ma. At this stage, the foreland basin sediments from Pakistan to India show significant supply from volcanic arcs and ophiolites of the Indus Suture Zone, indicating the absence of significant relief along the proto-Himalayan belt. Inversion of motion may have occurred within only 5 to 10 Myr after the collision onset, as soon as thicker and buoyant Indian crust choked the subduction zone. The arrival of thick Indian crust within the convergent zone 50-45 Myr ago led to progressive stabilization of the India/Asia convergent rate and rapid stabilization of the Himalayan shortening rate of about  $2 \text{ cm.yr}^{-1}$ . This first period also corresponds to the onset of terrestrial detrital sedimentation within the Indus Suture zone and to the Barrovian metamorphism on the Indian side of the collision zone. Equilibrium of the Himalayan thrust belt in terms of amount of shortening vs amount of erosion and thermal stabilization less than 10 Myr after the initial India/Asia contact is defined as the collisional regime. In contrast, the first 5 to 10 Myr corresponds to the transition from oceanic subduction to continental collision, characterized by a marked decrease of the shortening rate, onset of aerial topography, and progressive heating of the convergent zone. This period is

defined as the continental subduction phase, accommodating more than 30% of the total Himalayan shortening.

**Keywords:** Himalaya, Tectonics, Metamorphism, Sedimentology, Paleomagnetism, Subduction, Collision.

## **1. Introduction**

[2] The India-Asia collision, which gave rise to the Himalaya has been one of the most prominent geologic events in the Cenozoic. This collision had a profound impact on global climates and environments, greatly affecting atmospheric and oceanic circulations and floral to faunal distribution [Jaeger et al., 1989]. It probably also changed the asthenospheric circulation as recent tomography data argue for a deep subduction of the Indian lithosphere down to the transition zone [Van der Voo et al., 1999]. These authors suggest that between 1000 and 1500 km of Indian lithosphere have been subducted since the onset of India-Asia contact. In contrast, reconstruction of the initial geometry of the Indian crust shows that ~ 670 km of shortening have been accommodated at the scale of the Himalayan belt [e.g. DeCelles et al., 2002]. This suggests that the Himalayan shortening has been underestimated or that a part of the Indian upper crust has been early subducted with the rest of the Indian lithosphere. The occurrence of Early Eocene eclogites in the NW Himalaya [Pognante et Spencer, 1991; Guillot et al., 1997] clearly suggests that the distal part of the Indian continental margin was subducted and consequently the amount of shortening estimated by surface reconstruction is underestimated.

[3] The first aim of this paper is to give an overview of what occurred before and during the initial India-Asia contact from the Upper Cretaceous to the Early Eocene. The analysis of sedimentary processes, tectono-metamorphic processes, and paleomagnetic data at the end of

this period, allow us to define the concept of continental subduction. Then, the comparison of the stratigraphic record of the Eocene to Miocene foreland basins, the thermal evolution of the Indian crustal slices involved during the collision, and the amount of shortening accommodated within the growing orogen will be analyzed to discuss the concept of continental collision.

## **2. The Northwestern Himalayan belt.**

[4] In NW Himalaya, the complete evolution of the Himalayan belt from the Upper Cretaceous to the present-day is well preserved in both shallow (sedimentary) and deep (metamorphic) structural levels. The Himalaya rises abruptly from the Indo-Gangetic plain to high mountain peaks south of the Indus suture zone. The main tectonic units can be followed continuously in Western India from the south to the north (Figure 1). The north-dipping Main Frontal thrust (MFT) places the sub-Himalayan molasse belt over underformed Indo-Gangetic foreland basin sediments. The Main Boundary Thrust (MBT), active since at least 10 Ma [Burbank et al., 1996], places the Lesser Himalaya over the sub-Himalayan molasse. The Lesser Himalaya includes a 10 to 15 km thick section of Precambrian metasediments metamorphosed from low-grade to amphibolite facies metamorphic conditions [Le Fort, 1989]. The north-dipping Main Central Thrust (MCT), active since about 25-20 Ma [e.g., Hodges et al., 1996], places the 10 to 15 km thick Higher Himalayan Crystallines (HHC) over the Lesser Himalaya (Figure 2). The HHC comprise Precambrian basement and Paleozoic cover rocks metamorphosed during the Oligo-Miocene and intruded by Miocene leucogranites [Searle et al., 1992]. In western India, the Zaskar Shear Zone (ZSZ; western continuation of the South Tibetan Detachment system; Herren, 1987], separates the HHC from Late Precambrian to Eocene sediments of the Tethys Himalaya (Figure 1). This latter represents the shelf facies along the northern Indian continental margin [Gaetani and Garzanti, 1991], onto which the Spontang ophiolite was emplaced. North of the MCT, the Tethys

Himalaya is invariably detached at its base, suggesting an early decoupling of the sedimentary cover from its basement (Figure 4 in Guillot et al., 2000). The Precambrian to Orovician Haimanta formation crops out below the Tethys Himalayan cover and is generally considered as the northward prolongation of the HHC [Vannay and Grasemann, 1998] (Figure 2). However, Chawla et al. [2000] showed that this metamorphic unit is cross-cut by an underformed Ordovician granite suggesting that this unit is preserved from the Himalayan Tertiary metamorphism. The Tso Morari dome, that recorded the initial subduction of the Indian margin at  $55 \pm 7$  Ma [de Sigoyer et al., 2000], is sandwiched within the low-grade to amphibolitic Paleozoic metasediments. Finally, the Indus Suture Zone, squeezed between the Tso Morari dome and the Ladakh arc batholith, includes Indian continental slope and rise sediments (Lamayuru Formation; Bassoullet et al., 1983), slices of ophiolite mélange and Cretaceous blueschists together with Cretaceous to Paleogene forearc basin sediments (Figure 1) [Garzanti and Van Haver, 1988].

### **3. Tethys during the Cretaceous**

[5] The northward motion of the Indian Plate since the mid-Lower Cretaceous was responsible for the progressive closure of the Neotethys Ocean [Dercourt et al., 1993]. The Indus Suture Zone in western Ladakh includes two Tethyan paleo-subduction zones, beneath the Asian active margin and in a north-dipping intra-oceanic settings respectively [Reuber et al., 1987; Corfield et al., 1999; Mahéo et al., 2000; Roberston et al., 2000]. The subduction beneath the Asian active margin is well documented by the Dras calc-alkaline arc and the Ladakh-Kohistan batholith. This active margin represents a part of the greater Trans-Himalayan arc extending from Makran to southern Tibet [Beck et al., 1996]. All along this active continental margin, magmatism started synchronously at around 100-110 Ma [Debon et al., 1986]. The intra-oceanic subduction zone also observed from western Pakistan to southern

Tibet was active between 110 and 130 Ma [Beck et al., 1996; Gnos et al., 1997; Aitchison et al., 2000].

[6] The ophiolite obduction on the eastern part of the Tethys remains an unsolved complex problem. Emplacement of the Pakistan ophiolites took place either during Late Cretaceous time (ca 85 Ma; Beck et al., 1996) or at the Paleocene-early Eocene boundary (ca 65 to 50 Ma; Qayyum et al., 2001). The former interpretation assumes coupling between India and Africa-Arabia [e.g., Beck et al., 1996], although the two continental masses had been separated since mid-Jurassic times [e.g., Norton and Sclater, 1979; Molnar et al., 1988]. The latter relies on the occurrence of late Maastrichtian fossils in tectonic mélanges beneath the ophiolites, unconformably overlapped by lower Eocene sediments [Allemann, 1979].

[7] We assert that the Late Cretaceous ophiolite obduction took place either onto the Western Pakistan margin by transform movements caused by more rapid northward drift of the Indian Plate with respect to the adjacent Arabian Plate (Fig. 3), or in intra-oceanic settings far to the north, possibly in proximity to the Transhimalayan subduction zone. This hypothesis is supported by the docking of the Kohistan arc to the south Karakorum margin (south Asian margin) during the Upper Cretaceous [Treloar et al., 1996]. Similarly, in the Ladakh-Zaskar area (NW India), slices of arc-related lavas probably coming from the Spongtang-Nidar ophiolite and metamorphosed under blueschist conditions have been recently described in the Sapi-Shergol mélange [Mahéo, unpublished data]. As the Sapi-Shergol mélange corresponds to an accretionary wedge developed during the Upper Cretaceous in front of the south Asian subduction zone [Roberston et al., 2000], this suggests that the intra-oceanic arc represented by the Spongtang-Nidar ophiolite was incorporated into the Asian margin during Late Cretaceous time. Moreover, the stratigraphic record of the Tethys Himalaya indicates that the early obduction event did not directly involve the India passive continental margin during the Cretaceous. The major tectono-eustatic transgressive episode took place during mid-

Cretaceous time (late Albian to early Turonian; 98 to 91 Ma), at the end of rift-related volcanism and final detachment of India from Gondwana [Garzanti, 1999]. Pelagic oozes were next deposited in constant to gradually decreasing water depths until the early Maastrichtian, when a thick, upward-shallowing marly to carbonate succession accumulated [Nicora et al., 1987]. This latter stratigraphic unit has been given the inappropriate name «Kangi La Flysch» in the earlier geological survey reconnaissance studies [e.g. Kelemen and Sonnenfeld, 1983]. This name has led some authors to suggest that this “flysch” is related to an early obduction of the Spongtag Ophiolite onto the distal Indian margin [Searle, 1983; Searle et al., 1987].

[8] In contrast, Gaetani and Garzanti [1991] and Premoli Silva et al. [1991] showed that the Maastrichtian “Kangi La Formation” documents (1) progressive shallowing - rather than a «very rapid deepening event» [Searle, 1983] - in a mixed terrigenous/carbonate ramp setting, (2) approximately constant - rather than rapidly increased [Searle et al., 1987] - tectonic subsidence rates and (3) includes bioclasts, mud and quartzose silt to fine-grained sand derived from the Indian craton to the south (Figure 4). The absence of flexural tectonic subsidence and the lack of ophiolitic detritus indicate that these passive margin sediments were definitely not deposited in front of an obducting ophiolite [Kelemen et al., 1988].

[9] Thus, we propose that the Late Cretaceous ophiolite collage event may have taken place in an intra-oceanic setting far to the north of Greater India, possibly in the vicinity of the Transhimalayan subduction zone. The Spongtag ophiolite may have been offscraped during the Cretaceous stage of intra-oceanic subduction either within or at the northern side of Neotethys, and incorporated into the Asian accretionary prism after  $88 \pm 5$  Ma (age of andesitic arc sequence overlying the Spontang ophiolite, Pedersen et al., 2001). A possible age for the collision of the Ladakh-Kohistan intra-oceanic arc with the Asian margin at 65 Ma could be documented by a change in the velocity of the northward drift of India [Klootwijk et



al., 1992]. The final emplacement of the Spontang ophiolite onto the outer Zaskar shelf (Indian margin) occurred after the Early Eocene deposition of the Kong Formation [Garzanti et al., 1987].

#### **4. Timing of India-Asia contact**

[10] The age of the collision onset is extensively debated. In the western Himalaya, the proposed ages range from 65 to 45. The discrepancies between the inferred ages result from the use of different approaches and has consequences on the definition of continent-continent collision. Therefore, the sequence of the early orogenic events, from the first compressional deformation related to the initial subduction of the thin Indian margin, to final docking and rapid rise of the Himalayan range is still poorly understood. In the present paper, initial India-Asia contact is defined as the time when the edge of the Indian continent first arrived at the Kohistan-Transhimalaya trench, leading to the complete consumption of the Neotethys lithosphere, followed by continental subduction.

[11] The first India-Asia contact has been proposed at 65 Ma, based on significant lithospheric plate reorganization in the Indian ocean [Courtillot et al., 1986], evidence for the first India-Asia faunal exchange [Jaeger et al., 1989] and possible initiation of deformation in the Indian margin with emplacement of a tectonic mélange [Searle, 1983; 1987; Beck et al., 1996]. Paleomagnetic data from the Indian oceanic floor also record variations in the direction and in the velocity of the Indian plate motion at about 65 Ma [Klootwijk et al., 1992]. In the same way, the continental contamination described in the magmatism of the south Asian margin was related to the subducted Indian margin at about 60 Ma in Ladakh [Searle et al., 1987]. However, at this time, the Indian continent was located at about 1500 to 2000 km south of the Asian margin [Besse and Courtillot, 1988], which precludes an initial India-Asia contact at this time. Nevertheless, an initial Cretaceous-Paleocene Indian collision with the

Kabul block on the western part and the Kohistan-Ladakh arc is possible [Treloar and Coward, 1991]. Such an Early Paleocene contact could explain the onset of terrestrial fauna exchange between Asia and India at that time [Jaeger et al., 1989].

[12] In contrast, Dewey et al. [1989] and Le Pichon et al. [1992] suggest that the continental collision occurred during the Middle Eocene, at about 45 Ma. Such an age for the initial collision is incompatible with both the early foreland-basin stratigraphic record from Pakistan to Nepal, dominated by early Eocene detrital sediments from volcanic arc and ophiolites [Critelli and Garzanti, 1994; DeCelles et al., 1998; Najman and Garzanti, 2000] and with ultra high pressure metamorphism recorded at 55-45 Ma in the leading edge of the Indian plate. Klootwijk et al. [1992] and Acton [1999] correlated the sharp slowdown of spreading at 55+ Ma documented by magnetic anomalies on the Indian ocean floor with the true initial contact between India and Asia in the NW Himalaya. According to Besse et al. [1984], this was followed by a progressive eastward suturing between 50 and 40 Ma. Treloar and Coward [1991] and Rowley [1996], according to sedimentologic constraints, also argued that the collision first occurred in the western syntaxis at about 55-50 Ma and then at about 50-45 Ma in the central and eastern part of the range.

[13] A continuous record of transition from passive margin to collisional basin sedimentation is documented by accurately dated stratigraphic sections from the Higher Himalaya (Figure 4) [e.g., Baud et al., 1985; Nicora et al., 1987]. The major Late Paleocene (ca  $55 \pm 1$  Ma, conversion to Ma according to the Berggren et al., 1995 time scale) shallowing event in the distal part of the Indian margin was marked by an abrupt transition to peritidal dolostones. It was interpreted as flexural uplift related to initial contact of India and Asia [Garzanti et al., 1987]. The occurrence of debris flow conglomerates with limestone pebbles ranging from the Late Cretaceous to the Late Paleocene [Fuchs, 1987] also indicates that the onset of deformation is very close to the Paleocene/Eocene boundary (Figure 4).

[14] The sedimentary records of the Tethys Himalaya passive margin and the Transhimalayan active margin bear nothing in common from Cretaceous to Paleocene times, but begin to be closely comparable since the beginning of the Eocene, documenting the final closure of the Tethys between 55 and 50 Ma. Indeed, shallow-marine limestones yielding the Earliest Eocene nummulites (P6 to P7 zones, ca 54 to 51 Ma) are found on both sides of the Indus suture, from the distal Indian margin to the Transhimalaya forearc basin [Baud et al., 1985]. In Pakistan and in India, Early Eocene marine sediments were replaced by continental redbeds [Garzanti et al., 1987; Garzanti and Van Haver, 1988]. This continental redbeds unconformably overlying the Indian passive margin is characterized by the abrupt appearance of ophiolitic and volcanic detritus followed by arkosic detritus from the dissected roots of the Transhimalayan arc in fanglomerates capping the Indus forearc basin succession. This detrital sedimentation suggests the final closure of Tethys and active uplift of the Transhimalaya arc-trench system and ophiolitic rocks of the Indus suture by 50 Ma (Figure 4) [Garzanti et al., 1996]. As the initial Late Paleocene-Early Eocene contact took place at a low latitude of ca 8°N, the closure of the Neotethys determined an abrupt shift towards more arid climates, as documented by local evaporites and caliche paleosoils in continental redbeds [Garzanti et al., 1987].

## **5. The early collisional evolution**

[15] During the Early Eocene, suture-derived detritus replaced quartzose detritus fed from the Indian continent in the south during the whole Mesozoic and Paleocene (figure 4) [Nicora et al., 1987; DeCelles et al., 1998]. Stratigraphic dating of such a marked petrographic change represents the most accurate and reliable direct way to establish the precise age of final closure of the Neotethys, and to document its possible diachroneities in various segments of the future Himalaya [Rowley, 1996]. Farther to the south, the Eocene-Oligocene clastic

sediments observed on the Owen Ridge and presumed to represent the lower part of the Indus Fan include detrital K-feldspars with lead isotopic signatures pointing to an arc source in the Indus Suture Zone [Clift et al., 2000]. All provenance information thus consistently suggests that relief existed only on the northern side of the Indus suture zone in NW Himalaya in the earliest collisional stages. In contrast, the provenance data from Middle Eocene in Nepal indicate that Tethyan rocks were probably exposed and holding up relief south of the suture zone by that time [DeCelles et al., 1998; Robinson et al., 2001].

[16] The Early Thanetian to Lower Eocene units are capped by extensive lateritic paleosols throughout northern Pakistan, India and Nepal, documenting prolonged exposure of sediments and lack of significant subsidence during most of the late Eocene to the Oligocene [Pivnik and Wells, 1996; DeCelles et al., 1998]. Such a long stage of negligible sediment accumulation - from about 50 Ma to 25 Ma - has been related either to a transition from the low-strength collision (subduction of the thinned Indian continental-margin crust) to high-strength collision (underthrusting within the collision zone zone by unstretched Indian crust; Najman and Garzanti 2000) or to southward migration of the flexural bulge [DeCelles et al., 1998]. If the latter hypothesis is correct, the underlying syn-collisional sediments would represent back-bulge deposits. In addition, most of the sediment volume deposited in the main foredeep depozone of the foreland basin system would not be preserved anywhere along the Himalayan range, excepting perhaps the limited outcrops of Lower Eocene redbeds in the Tethys Himalaya of Zaskar and southern Tibet [Garzanti et al., 1987; Willems et al., 1996]. A huge volume of clastic sediments was inferred to have been derived from the rising Himalaya and deposited from the Late Eocene to the Early Miocene in remnant-ocean basins from Katawaz to Makran. This deltaic to turbidite system might have represented the major depocenter of orogenic sediments derived from the Himalayan uplands at a stage of general bypassing and westward axial transport [Qayyum et al., 2001].

[17] The Balakot redbeds widely exposed in the Hazara re-entrant, was previously thought to represent an up to 8 km-thick continuous stratigraphic succession of the Early to Middle Eocene age [Bossart and Ottiger, 1989]. Recently Najman et al. [2001] showed that the Balakot formation contains micas with Oligocene detrital Ar/Ar ages. There is thus no apparent exception to the limited thickness of the early collisional foreland basin sediments along the Himalaya, documenting a very low subsidence related to a low flexural bulge.

[18] Since the latest Oligocene, renewed foreland basin subsidence ( $> 0.2 \text{ mm.yr}^{-1}$ ) was associated with thrusting and accretion of the Himalayan orogenic wedge and characterized by thick, fine-grained terrigenous successions with metamorphic grade steadily increasing in time from a very low grade (pre-Himalayan) to a low grade clastic sediments (Himalayan overprint). The youngest metamorphic imprint is revealed by the wealth of slate to phyllite lithics suddenly supplied to the Pakistan and Indian foreland region [Critelli and Garzanti, 1994; Najman and Garzanti, 2000], which yield detrital micas with metamorphic Ar/Ar ages of 36-40 Ma [Najman et al., 2001] and 28-25 Ma [Najman et al., 1997], respectively (figure 4). Metamorphic grade of lithic fragments reached the garnet zone. These garnet-bearing micaschists were tectonically exhumed and eroded around 22 Ma, as indicated by the cooling ages of detrital micas within the Kasauli and Dharamsala Formations [Najman et al., 1997; Najman and Garzanti, 2000]. These events may record motion along the Main Central thrust, with unroofing of Himalayan metamorphic rocks and active till the deposition of the Lower Dharamsala Formation at 17 Ma (Figure 4) [White et al., 2000].

## **6. Timing of the Early Himalayan metamorphism**

[19] In NW Pakistan, two distinct phases of the early metamorphism are distinguished (Figure 5). The eclogitic event ( $> 25 \text{ kbar}$ ,  $> 600^\circ\text{C}$ ) recorded in the Kaghan nappe [Pognante and Spencer, 1991; O'Brien, et al., 2001] and also in the partially preserved eastern part of the

Nanga Parbat syntaxis [Le Fort et al., 1997] is related to the early subduction of the Indian Plate below the Kohistan-Ladakh arc. The HP granulitic facies metamorphism ( $13 \pm 3$  kbar,  $750 \pm 50^\circ\text{C}$ ) associated with partial melting is related to the thickening of the Indian plate [Treloar et al., 1989; Pognante et al., 1993]. The spatial distinction between the eclogitic and the granulitic units is difficult because they are invariably imbricated within a thrust pile [Treloar et al., 1989]. Tonarini et al. [1993] dated The eclogitic assemblage of Kaghan is dated between between 49 and 46 Ma by Sm/and, U/Pb and SHRIMP methods [Tonarini et al., 1993; Spencer and Gebauer, 1996; Kaneko et al. 2001]. The upper structural levels of the Nanga Parbat massif and Kaghan upper rocks were buried to pressure of c. 10 kbar and heated to temperature of c.  $650^\circ\text{C}$  at 46-41 Ma [Smith et al., 1994; Zeitler and Chamberlain, 1991; Foster et al., 2002]. Finally, Treloar and Rex [1990], Chamberlain et al. [1991] and Tonarini et al. [1993] showed by Ar/Ar, Rb/Sr and U/Pb thermochronology that cooling below  $500^\circ\text{C}$  was completed 40 Myr ago (Figure 5). Similar to the Pakistan, two distinct phases of the early metamorphism are temporal and spatially distinguished in the NW India. The eclogitic event ( $> 25$  kbar,  $600 \pm 50^\circ\text{C}$ ) is recorded south of the Indus suture zone in the Tso Moriri dome (Figure 1) [de Sigoyer et al., 1997; Guillot et al., 1997, O'Brien et al., 2001]. The eclogitic event is dated at  $55 \pm 6$  Ma by U/Pb, Lu/Hf and Sm/Nd methods [de Sigoyer et al., 2000]. This metamorphic age is interpreted as the initial age of Indian continental subduction beneath the Asian margin [de Sigoyer et al., 2000]. The retrogression under amphibolitic facies metamorphic conditions occurred at  $47 \pm 2$  Ma according to the Sm/Nd, Rb/Sr and Ar/Ar datings (Figure 5). Moreover, apatite fission track ages of  $46 \pm 2$  Ma from the Ladakh intrusives in the Kargil area suggest that the area was not affected by any post-Middle Eocene thermal events [Lal and Nagpaul, 1975]. Clift et al. [2002] also report apatite fission track ages ranging between 44 and 27 Ma, and Ar/Ar biotite ages of 49-44 Ma for the Ladakh Batholith.

[20] The low geothermal gradient preserved in the eclogitic units of Kaghan and Tso Morari is clearly related to the subduction of the Indian margin below the Kohistan-Ladakh arc between 55 and 50 Ma. In contrast, the first Barrovian metamorphism (HP amphibolitic to granulitic facies conditions) recorded both by the partly exhumed eclogitic unit and by the granulitic unit located close to the suture zone during the Middle Eocene ( $\approx 45$  Ma) is more difficult to explain. This event occurred less than 10 Myr after the initial impingement of India against Asia, while the classical conductive thermal model suggests that a minimum of 20 to 30 Myr is necessary for a previously thickened crust to relax thermally [England and Thompson, 1984].

[21] In NW Himalaya, the first granulitic metamorphic event and the associated crustal melting (50-40 Ma) followed immediately the eclogitic metamorphic event (55-45 Ma). This suggests that these tectono-metamorphic events could be intimately related and related to the breakoff of the subducted India slab during Early to Middle Eocene time [Guillot et al., 1997; Chemenda et al., 2000, Kohn et al., 2002]. This hypothesis is supported by an important remelting of the Ladakh batholith between 50 and 46 Ma [Weinberg and Dunlap, 2000].

[22] Southward, another and younger metamorphic event ( $11 \pm 3$  kbar,  $700 \pm 50^\circ\text{C}$ ) is well preserved in the HHC from Zaskar to Bhutan [e.g. Guillot et al., 1999 for review]. In NW India, the HHC slab is tectonically separated from the Tso Morari dome by the Zaskar synclinorium (Figure 1) and clearly corresponds to a distinct unit (Figure 2). The Eo-Himalayan metamorphic event recorded by the HHC ranges between 37 and 30 Ma [Searle et al., 1992; Prince et al., 1999]. Moreover, Vance and Mahar [1998] showed that the onset of prograde garnet growth in Zaskar started at about 33 Ma. In the same way, Prince et al. [1999] dated leucosomes from Gahrwal at about 37 Ma, suggesting that the HHC was thermally reequilibrated at this time (Figure 5).

[23] In order to explain the heating recorded by the HHC during the Oligo-Miocene over a short period of time, Guillot and Allemand [2002] have tested, by a two-dimensional thermal model, the time necessary to heat a 10 km thick continental slice underthrust below a high-heat producing zone that could represent the thickened internal Himalayan zone. Guillot and Allemand [2002] have shown that the best way to reproduce the P-T conditions recorded by the HHC is to impose a decoupling in depth of the HHC from the rest of the subducting Indian plate. Such decoupling allows the HHC both to remain at a constant depth and to be heated up. 10 million years were necessary to reach a temperature of 700°C in the HHC, while the underthrust Indian plate remained at a relatively low temperature (< 600°C). To preserve high temperature (> 600°C) during the exhumation of the HHC, a high vertical rate (> 3 mm.yr<sup>-1</sup>), similar to the present-day uplift rate, was required. This suggests that the MCT and STDS, the tectonic boundaries of the HHC were probably active over a short period of time (< 10 Myr) during the Early Miocene, compatible with the short period of leucogranite emplacement and geochronological records of the main MCT activity between 25 and 18 Ma [Hubbard and Harrison, 1989; Guillot et al., 1994; Hodges et al., 1996].

[24] In south Karakorum (NE Pakistan), the discovery of Neogene granulitic rocks associated with migmatites and numerous mantle-derived magmatic rocks (Karakorum Metamorphic Complex) in a setting of global north-south shortening (Rolland et al., 2001) strongly suggests that interaction between the thickened Asian crust and the underlying mantle occurred during the India-Asia convergence. Eastward, in southern Tibet, potassic Neogene magmatism have also been observed (Mahéo et al., 2002). The origin of the south Karakorum granulites and the associated Neogene magmatism all along the southern Tibet are discussed in light of a second slab breakoff of the subducting Indian slab, starting at about 25 Ma [Mahéo et al., 2002].

## **7. Modeling of the Himalayan shortening**



[25] According to the available paleomagnetic data, Dewey et al. [1989] and Le Pichon et al. [1992] estimated a total convergence of 2300-2150 km in the western syntaxis since 45 Ma, whereas Molnar and Tapponnier [1975] estimated a total convergence of  $3000 \pm 500$  km. By backward motion of Asian and Indian lithospheric blocks, Replumaz and Tapponnier [2003] also estimated 3000 km of convergence if the initial India-Asia contact is at 55 Ma.

[26] The Himalayan shortening (south of the Indus suture zone) is estimated by mass-balanced cross-sections at  $\sim 670$  km from Pakistan to Sikkim [e.g. DeCelles et al., 2002]. By plate reconstruction, Himalayan shortening is estimated at  $1250 \pm 250$  km [Achache et al., 1984; Besse et al., 1984; Powell et al., 1988; Dewey et al., 1989; Patzelt et al., 1986; Matte et al., 1997]. This difference is explained by the existence of a Greater India, extended up to 650-700 km, north of the present-day Indus suture zone, and consisting of all of the Indian lithosphere that has been subducted beneath southern Tibet [Klootwijk et al., 1979; Patriat and Achache, 1984; DeCelles et al., 2002]. The existence of the Greater India is compatible with the original fit of the North Indian margin with the North Australian margin at 160 Ma [Powell et al., 1988; Dercourt et al., 1993; Matte et al., 1997]. The existence of a Greater India has important consequences for the earlier evolution of the Himalayan belt and implies that a part of the Indian continental lithosphere was totally subducted before 45 Ma. Moreover, if the initial India-Asia contact occurred at 55 Ma rather than at 45 Ma, the amount of north-south India-Asia convergence is underestimated in the model of Dewey et al. [1989] and Le Pichon et al. [1992]. In figure 6, we propose to reestimate by balanced crustal cross-section the amount of Himalayan shortening through time taking into account three new facts presented in this manuscript. As discussed above, we assumed first that the initial impingement of India against Asia was at 55 Ma. Second, we took into account the occurrence of Himalayan eclogites showing the subduction of the Indian margin below the southern Tibet to a minimum depth of 100 km. Third, we considered that the Haimanta

formation is distinct from the HHC. By balanced cross-section, we estimated a minimum Himalayan shortening of about 400 km between 55 and 40 Ma and a total Himalayan shortening of about 1100 km between 55 Ma and the present-day (Figure 6). These values are largely superior than the previous shortening estimates because it takes into account the earlier subduction of the Indian plate north of the Indus suture zone, and are in the same order as paleomagnetic reconstruction estimates.

[27] In order to independently constrain the amount of convergence accommodated by the Indian plate since the initial India-Asia contact, we used an original method. In the following, the Himalayan shortening is defined as the displacement of India relative to the Indus suture (which itself could be moving). Therefore, the India-Asia convergence is defined as the sum of the Himalayan shortening plus Asian contraction including the lateral extrusion.

[28] In order to interpolate the convergence velocity data at different periods of time, the continental subduction of India beneath Asia is modeled as a two converging block system (figure 7). In this one-dimensional model, we consider the forces per unit length along the x axis which are applied to India: the force which drives the convergence ( $D$ , unit [Pa/m]), and a force ( $f$  unit [Pa/m]) which resists to it (figure 7). The equation of the movement is:

$$D + f = \rho * L * H \frac{dv}{dt} \quad (1)$$

where  $v$  is the velocity of convergence,  $dv/dt$  its time derivative and  $L*H$  is the sectional area of the system subject to deformation. The force which resists to the collision which comes from various effects (viscous shear, folding, sliding along faults etc...) is approximated by the shear stress ( $\tau$ ) due to the change of velocity in the vertical direction in a viscous flow:

$$\tau = \frac{\mu}{2} \frac{dv}{dz} \quad (2)$$

where  $\mu$  is the apparent viscosity of the material. Equation (1) writes:

$$D + L\tau = \rho LH \frac{dv}{dt} \quad (3)$$

The derivative of the velocity along the vertical direction scales as  $v/H$  where  $v$  is the convergence velocity. The convergence between India and Asia extended from 55 Ma to present time. We assume that the thickness and the length of the zone subject to deformation and the driving force are constant over this period of time. Equation (1) becomes:

$$\frac{dv}{dt} = \alpha v + \beta \quad (4)$$

where  $\alpha$  and  $\beta$  are two constants depending on  $L, H, D, \rho$  and on the velocity profile in the vertical direction. The equation (4) can be solved and we obtain an exponential law for the convergent rate:

$$v = a \exp\left[\frac{b - T}{c}\right] + d \quad (5)$$

where  $T$  is the time (0 is present time, past time is negative). The constants  $a$  and  $d$  have the dimension of a velocity [ $\text{ms}^{-1}$ ] and  $c$  has the dimension of time;  $a$  represents the velocity at  $T = b$ ,  $d$  is the asymptotic velocity,  $c$  controls the rate of decrease of the velocity through time. The parameters of this exponential law are computed from data of the convergent velocity by using a least-squares method. Although, this model has very strong assumptions (constant driving force and length scale for the deformed area, viscous deformation as the resisting force to the convergence), we think that it provides the best simple analytical form in order to interpolate the velocities of convergence.

[29] According to the available paleomagnetic data [Patriat and Achache, 1984; Courtillot et al., 1986; Besse and Courtillot, 1988; Dewey et al., 1989; Klootwijk et al., 1992; De Mets et al., 1990; Acton, 1999], reconstructions of the motion of India with respect to Eurasia allow to distinguish 4 periods since 65 Ma:

- 1- anomaly 30 (~65 Ma) to anomaly 24 (55 Ma): very fast convergence of about  $18 \pm 5 \text{ cm.yr}^{-1}$ ;
- 2- anomaly 24 (~55 Ma) to anomalies 22-21 (51-49 Ma) : sharp slowdown from  $18 \pm 5 \text{ cm.yr}^{-1}$  to  $10 \pm 2 \text{ cm.yr}^{-1}$ ;
- 3- anomalies 22-21 (51-49 Ma) to anomaly 18 (43 Ma) : progressive decrease down to  $6.0 \pm 1 \text{ cm.yr}^{-1}$ ;
- 4- reorganization of Indian ocean spreading leading to the  $4.5 \pm 0.5 \text{ cm.yr}^{-1}$  convergent velocity from 20 Ma until present.

[30] According to these data, we constructed a curve for the India/Asia convergence (Figure 8). The fitting of these data allow us to estimate numerically the parameters a,b,c,d (Table 1).

	a	b	c	d
Mean	$3.76 \text{ cm.yr}^{-1}$	4.7 Myr	$0.180 \text{ Myr}^{-1}$	$4.5 \text{ cm.yr}^{-1}$
Fast	$0.99 \text{ cm.yr}^{-1}$	3.8 Myr	$0.178 \text{ Myr}^{-1}$	$5.0 \text{ cm.yr}^{-1}$
Slow	$5.58 \text{ cm.yr}^{-1}$	5.2 Myr	$0.184 \text{ Myr}^{-1}$	$4.0 \text{ cm.yr}^{-1}$

**Table 1:** Computed values of the parameters used to determine the India-Asia convergence and the associated velocity.

Then, the integration of these parameters within equation 3 allow a numerical estimate of the amount of north-south convergence recorded by the India/Asia suture and the correspondent rate, for different selected periods (Table 2). The uncertainties are quoted at  $1\sigma$ .

Period	Total convergence	Rates
55 Ma -> 50 Ma	$670 \pm 166 \text{ km}$	$13.4 \pm 3.3 \text{ cm.yr}^{-1}$
50 Ma -> 40 Ma	$703 \pm 116 \text{ km}$	$7.0 \pm 1.2 \text{ cm.yr}^{-1}$
40 Ma -> 20 Ma	$945 \pm 110 \text{ km}$	$4.7 \pm 0.6 \text{ cm.yr}^{-1}$
20 Ma -> 0 Ma	$897 \pm 103 \text{ km}$	$4.5 \pm 0.5 \text{ cm.yr}^{-1}$

55 Ma -> 0 Ma	$3215 \pm 496$ km	$5.8 \pm 0.9$ cm.yr <sup>-1</sup>
45 Ma -> 0 Ma	$2140 \pm 271$ km	$4.7 \pm 0.6$ cm.yr <sup>-1</sup>

**Table 2:** North-south India/Asia convergence and associated rates computed with the data of the Table 1.

[31] If the initial India-Asia contact occurred at 45 Ma, we estimate a total India-Asia convergence of  $2140 \pm 271$  km, similar to the Dewey et al. [1989] and Le Pichon et al. [1992] estimates. Similarly, the  $3215 \pm 496$  km of convergence estimated since 55 Ma (Table 3) is in the same order with the  $3000 \pm 500$  km of convergence calculated by Molnar and Tapponnier [1975]. As we demonstrate that the initial India-Asia contact occurred at 55 Ma rather than 45 Ma (see discussion above), an excess of 1000 km of convergence has been accommodated during the first 10 Myr and confirms the existence of the Greater India. It is also noticeable that the inflexion point of the velocity curves is located between 50 and 43 Ma in our model, i.e. very close to the supposed timing of collision at 45 Ma proposed by Dewey et al. [1989] and Le Pichon et al. [1992]. Thus, our preliminary conclusion is that the collision classically defined in the literature at 45 Ma corresponds in our modeling to the stabilization of the India/Asia convergent rate, 10 million years after the initial contact.

[32] In order to calculate the amount of shortening recorded only in the Himalayan belt (the Indian side of the convergent zone) by the equation (3), we fixed the following boundary conditions. We first assumed that during the initial India-Asia contact at 55 Ma, the Himalayan shortening rate is equal to the velocity of the Indian plate ( $18 \pm 5$  cm.yr<sup>-1</sup>). We also fixed the present-day convergent rate within the Himalaya belt at 2 cm.yr<sup>-1</sup> [Bilham et al., 1997]. Second, in order to estimate the Himalayan shortening rate per selected periods, we impose a total Himalayan shortening ranging between 1100 and 1600 km (Figure 9) deduced from the initial geometry of the Greater India (Figure 3) [Dalziel et al., 1987; Dercourt et al., 1993; DeCelles and DeCelles, 2001] and tomographic data [Matte et al., 1997; Van der Voo

et al., 1999]. The difference between the total India-Asia convergence (Table 2) and the calculated Himalayan shortening (Table 4) corresponds to the total amount of shortening accommodated by the Asian plate (Table 5).

	a	b	c	d
Mean	0.82 cm.yr <sup>-1</sup>	5.0 Myr	0.6 Myr <sup>-1</sup>	2.0 cm.yr <sup>-1</sup>

**Table 3:** Computed values of the parameters used to determine the Himalayan shortening and the associated rate.

Period	Convergence	Velocity
55 Ma -> 50 Ma	345 ± 140 km	6.9 ± 2.9 cm.yr <sup>-1</sup>
50 Ma -> 40 Ma	210 ± 70 km	2.1 ± 0.7 cm.yr <sup>-1</sup>
40 Ma -> 20 Ma	400 ± 20 km	2.0 ± 0.1 cm.yr <sup>-1</sup>
20 Ma -> 0 Ma	400 ± 20 km	2.0 ± 0.1 cm.yr <sup>-1</sup>
55 Ma -> 0 Ma	1355 ± 250 km	2.4 ± 0.4 cm.yr <sup>-1</sup>

**Table 4:** Himalayan shortening and associated velocities per selected periods.

[33] By using equation (3), we estimated that the only way for the initial boundary conditions to converge towards a unique numerical solution is that the Himalayan shortening velocity decrease dramatically to 2.1 cm.yr<sup>-1</sup> between 55 and 40 Ma. Moreover, if we use the 1100 km value of the Himalayan shortening, which is regard as the most realistic (DeCelles et al., 2002), the deceleration must happen between 55 and 51 Ma. The fact that the estimated shortening velocity in the Himalaya was already close to the present-day velocity as early as 50-45 Ma suggests that the convergent zone tends towards a steady state in terms of accretionary flux earlier than has classically been suggested (Hodges, 2000].

Period	Convergence	Velocity
55 Ma -> 50 Ma	$325 \pm 280$ km	$6.5 \pm 5.6$ cm.yr <sup>-1</sup>
50 Ma -> 40 Ma	$493 \pm 186$ km	$4.9 \pm 1.9$ cm.yr <sup>-1</sup>
40 Ma -> 20 Ma	$545 \pm 115$ km	$2.7 \pm 0.6$ cm.yr <sup>-1</sup>
20 Ma -> 0 Ma	$497 \pm 108$ km	$2.5 \pm 0.5$ cm.yr <sup>-1</sup>
55 Ma -> 0 Ma	$1860 \pm 689$ km	$3.4 \pm 1.25$ cm.yr <sup>-1</sup>

**Table 5:** Asian shortening and associated velocity.

[34] Our fitting suggests that 30% of the total Himalayan shortening ( $345 \pm 140$  km) was accommodated during the first 5 Myr. During this period, Himalayan shortening was greater than Asian shortening. Moreover, almost 50% of the total Himalayan shortening ( $555 \pm 210$  km) was accommodated during the first 15 Myr. This value is higher than the 400 km estimated by Patriat and Achache [1984] and also by mass-balance reconstruction. This suggests that on the one hand we have overestimated the total Himalayan shortening ( $1355 \pm 250$  km) and consequently, the total Himalayan shortening is closer to 1000 km rather than 1500 km as previously suggested. On the other hand, this result emphasizes the major role played by the first 5 to 15 million years for Himalayan evolution. As our estimate of the total Asian shortening ( $1860 \pm 746$  km) is similar to the  $1700 \pm 610$  km previously estimated [Tapponnier et al., 1986; Halim et al., 1988], we suggest that we did not overestimate the total Himalayan convergence.

## 10. Discussion

[35] The demonstration that the initial India-Asia contact took place close to 55 Ma rather than at 50 or 45 Ma is crucial to understand of the Himalayan building processes. Fitting of paleomagnetic data, and composition of the early foreland basin deposits derived mainly from arc and ophiolite rocks of the suture zone, show that during the first 5 Myr, a large part of the

thin Indian margin was subducted below the Asian margin without creating high relief on the Himalayan side (Figure 10A). Moreover, the occurrence of marine sediments up to the Late Ypresian (50 Ma) in the Indus suture zone suggests that this first stage occurred mainly below sea level. Our estimate of the amount of shortening shows that  $400 \pm 140$  km of Indian crust was subducted during this short period (Figure 10B). It is also noteworthy that during this period, the Himalayan shortening was greater than Asian shortening.

[36] The fact that only a small quantity of deep material is observable at the surface is compatible with the absence of erosion during this period. From a thermal point of view, the facts that the low-temperature eclogitic unit recorded isothermal decompression during this initial period (Figure 5) and that the HHC had not yet reached its maximum temperature also suggest that the thermal equilibrium was not reached. We can conclude that the first 5 Myr in the life of the Himalayan belt, after the initial India-Asia contact at about 55 Ma, represent a transitional period, corresponding to the complete subduction of the continental lithosphere. Only the Tethys sedimentary cover is decoupled from its basement that is subducted within the mantle wedge [Guillot et al., 2000] (Figure 10A, B). We define this stage as the continental subduction period.

[37] The 50-45 to 25-20 Ma interval corresponds to a major change in the India-Asia convergent rate, Himalayan shortening rate as well as in the sedimentary and metamorphic processes. During this period, the suturing was completed all along the belt [Klootwijk et al., 1992; Rowley, 1996], and  $610 \pm 90$  km of Indian margin was accreted within the Himalayan wedge (Figure 10C, D). The transition from the continental subduction period to this new one could correspond to the slab breakoff of the subducted Indian margin leading to the first amphibolitic to granulitic metamorphic event and to a low flexural bulge of the Indian plate. The HHC were progressively underthrust below the Tethys Himalaya (Figure 10B,C). However, this period is characterized by the apparent absence of subsidence, lack of



deposition, and widespread pedogenesis in the foreland basin. Convincing reasons for this have not been proposed so far. They may include: 1) uplift of the foreland basin due to docking of the subduction zone by arrival of the Indian crust with normal thickness represented by the HHC; 2) subsequent erosion of the main foreland basin depozone and preservation only of the fore bulge

Stage	Oceanic Subduction	Continental Subduction	Continental Collision	Steady-State Collision
Age	> 55 Ma	55-50 Ma	50-25 Ma	25-0 Ma
Crust type	oceanic	thin Indian crust	normal Indian crust	normal Indian crust
India-Asia velocity	18 cm.yr-1	18 -> 10 cm.yr-1	10 -> 4.5 cm.yr-1	4.5 cm.yr-1
India/Asia convergence	0 cm.yr-1	670 ± 166 km	1625 ± 220 km	1122 ± 105 km
Himalayan velocity	18 cm.yr-1	18 -> 2.5 cm.yr-1	2.5 -> 2.0 cm.yr-1	2.0 cm.yr-1
Himalayan shortening	0 km	345 km	510 ± 80 km	500 ± 25 km
Asian velocity	0 cm.yr-1	< 5 cm.yr-1	< 5.0 -> 3.0 cm.yr-1	2.6 cm.yr-1
Asian shortening	0 km	< 250 km	493 ± 186	622 ± 110
Subsidence rate	~ 0.02 mm.yr-1	~ 0.1 mm.yr-1	< 0.01 mm.yr-1	> 0.2 mm.yr-1
Sedimentary facies	shelf sedim	deltaic redbeds	unconformities	fluvial molasse
Sandstone petrography	quartzarenites	volc+ophio detritus	-----	metam detritus
Indian edge metamorphism	none	eclogite facies	amphibolite facies	greenschist facies
HHC metamorphism	none	none	granulite facies	anatexis
Asian Margin	arc volcanism	cont sedim	uplift	uplift

**Table 6:** From subduction to collision dynamics in Himalaya

zone closer to the stable foreland; 3) sediment bypass and/or shift of sediment depocenters elsewhere (Katawaz – Makran – early Indus Fan); 4) very low erosion rates as a result of either still very low mountain relief, or arid climate. It is noteworthy that the Andes for instance appear to be a highly-elevated mountain belt, but scarce detritus is accumulated in the arc-trench systems from Peru to Northern Chile due to pronounced aridity [Montgomery et al., 2001]. A marked shift is documented in the Tethys Himalayan succession from humid equatorial climates at late Paleocene times (pure quartzarenites with quartz pitted in lateritic paleosols) to more arid conditions in the Early Eocene (local evaporites, caliche paleosols in redbeds). This was apparently related to the closure of the Neotethys, and thus decrease source of humid air masses [Bossart and Ottiger, 1989]. However, why foreland basin

subsidence is everywhere in the front of the range from N Pakistan to Nepal very low until about 20 Ma remains a puzzling open question. Deep-water facies were apparently not deposited anywhere in the Himalayan foreland basin, which was always subaerial to very shallow-marine (e.g., base of foreland basin sequence directly alluvial or transitional marine at most, a different situation with respect to other orogens). Whatever the tectonic reason may be, this conclusively points to a very shallow continental subduction angle ( $< 10^\circ$ ).

[38] Progressive thickening of the radiogenic upper Indian crust allowed progressive warming of the internal Himalayan wedge responsible for the Eo-Himalayan metamorphism recorded in the HHC. During this period, the whole Asian plate started to be strongly affected by the India-Asia collision [e.g. Tapponnier et al., 1986; Lacassin et al., 1997; Replumaz and Tapponnier, 2003], suggesting that the stress induced by the India-Asia convergence was progressively transmitted to the hinterland part of the system and the width of the collision zone grew northward with time [Tapponnier et al., 2001]. Since 45 Ma, the velocity of northward indentation of the Indian plate became higher than the velocity of Indian plate subduction. This stage corresponds to the onset of the collision as defined by Dewey et al. [1989] and Le Pichon et al. [1992].

[39] Since about 25-20 Ma, progressive southward propagation of the thrust front and thrusting of the Lesser Himalaya below the MCT enhanced the progressive exhumation and erosion of the warm HHC wedge (Figure 10E). Such a rapid exhumation of the thermally relaxed HHC during the Miocene is responsible for their partial decompression melting all along the belt [Harrison et al., 1998; Guillot et al., 1999]. Deposition of thick sequences of alluvial sediments in the foreland basin since the Early Miocene [DeCelles et al., 1998; 2001; Najman and Garzanti, 2000; Najman et al., 2001] and throughout the Miocene [White et al., 2002] document markedly increased erosion and sedimentation rates [Burbank et al., 1996]. At the lithospheric scale, northward indentation velocity of the Indian plate ( $3 \text{ cm.yr}^{-1}$ )

became greater than its subduction velocity ( $2 \text{ cm.yr}^{-1}$ ), inducing its progressive steepening and roll over (Figure 10E). We propose that the underthrusting of the Indian crust could be a continuous process with a constant velocity of about  $2 \text{ cm.yr}^{-1}$  since 50-40 Ma. In contrast, the tectonic and thermal activities of the metamorphic units such as the HHC were probably a discontinuous processes with long periods of no tectonic activity ( $> 10 \text{ Myr}$ ) followed by rapid ( $\text{cm.yr}^{-1}$ ) and short tectonic activity ( $< 5 \text{ Myr}$ ) which was balanced by a strong erosion. We propose to define this period ranging between 50 Ma and the present-day as the continental collision period.

## 11. Conclusion

[40] Constraints from stratigraphy, paleomagnetism, geochronology and tectonophysics in the NW Himalaya indicate that the initial India-Asia contact took place very close to the Paleocene/Eocene boundary (55 Ma) after a long period of oceanic subduction. A major decrease in plate velocity (from  $18 \text{ cm.yr}^{-1}$  to  $10 \text{ cm.yr}^{-1}$ ) from  $55^+$  Ma to 50 Ma is interpreted as the effect of the India-Asia contact. Final closure of Neotethys was recorded by forced shoaling of marine sediments in Zaskar at the Paleocene/Eocene boundary ( $55 \pm 0.5 \text{ Ma}$ ), followed in the late Ypresian (50 Ma) by deposition of deltaic sediments derived from arc and ophiolite rocks incorporated in the obducted Asian accretionary prism. Petrography of foreland basin clastics indicates that marine seaways between India and Asia did not exist anymore at that time, and that major relief formed only on the Asian side of the suture, not along the proto-Himalayan belt. At the end of this period, a small part of the subducted Indian continental margin was exhumed at the base of the Indian crust. Subduction of the Indian plate probably ended by slab breakoff. It was followed by initial Himalayan-wedge thickening by underthrusting of continental units and emplacement of thin-skinned thrust-sheets.

[41] From 50-45 Ma to 25-20 Ma, the underthrusting of the HHC led to the thickening and warming of the Himalayan orogenic wedge, as reflected by the Eo-Himalayan metamorphism. Plate velocity decreased progressively down to 5-6 cm.yr<sup>-1</sup> since 45 Ma, and was stabilized thereafter, indicating that kinematics equilibrium was progressively reached. At the surface, this period corresponds to a very long stage of negligible sediment accumulation, suggesting virtually no foreland basin subsidence. Low erosion rates at this stage can be related either to a lack of significant relief in the proto-Himalayan belt related to a low-angle continental subduction plane and/or to arid climates at subtropical latitudes before the onset of the monsoon system.

[42] Since 25-20 Ma, the HHC wedge was exhumed along the MCT and the STDS, and alluvial clastic sediments were deposited in the thick foreland basin. The orogenic wedge rapidly grew in both width, as documented by southward propagation of the thrust front until the monsoonal system, profoundly altering atmospheric circulation patterns and earth's climates, was established and strengthened between ca 11 and 6 Ma [Quade et al., 1995].

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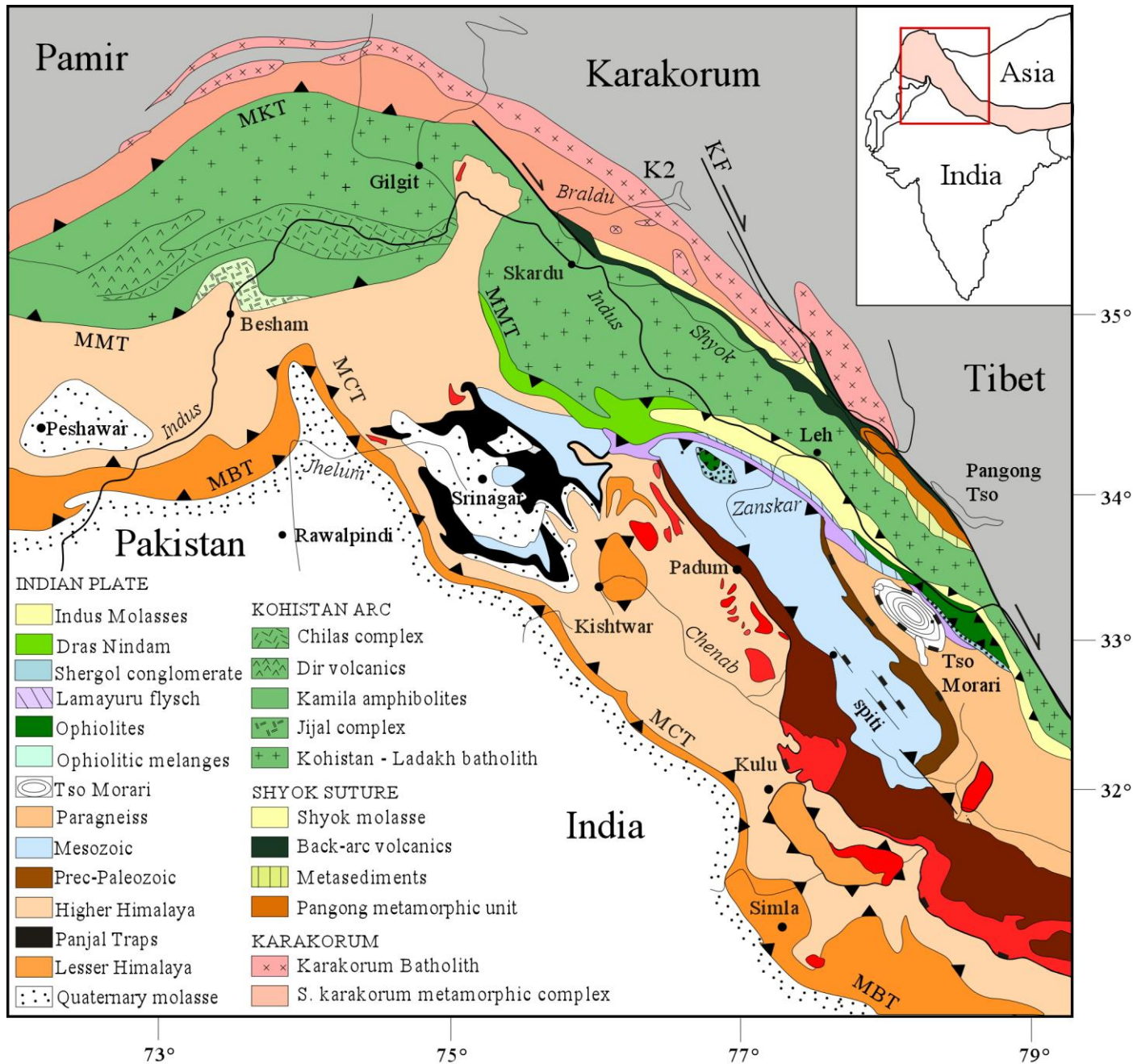


Figure 1 : Geological map of the Northwestern Himalaya with the location of the figure2. [after Chawla et al. 2000, Searle et al., 1999 ; Pêcher and Le Fort, 1999 and Mahéo et al., 2000]

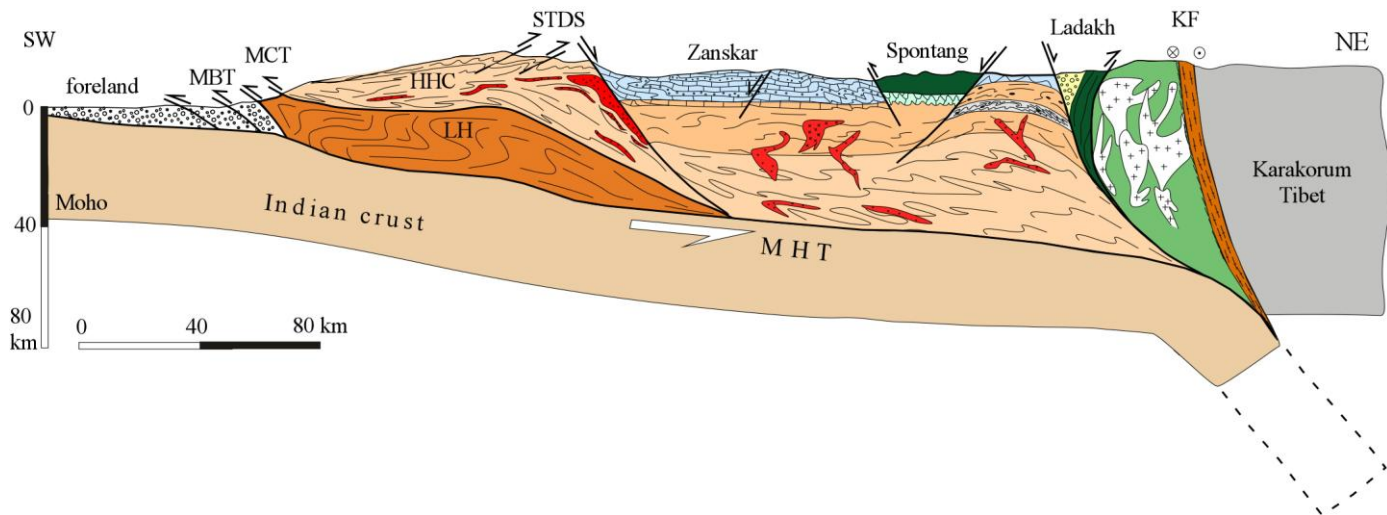


Figure 2 : Geological cross section of the Himalayan belt. The northward prolongation of the MHT is deduced from Ni and Barazangi [1984] and Nelson et al. [1996]

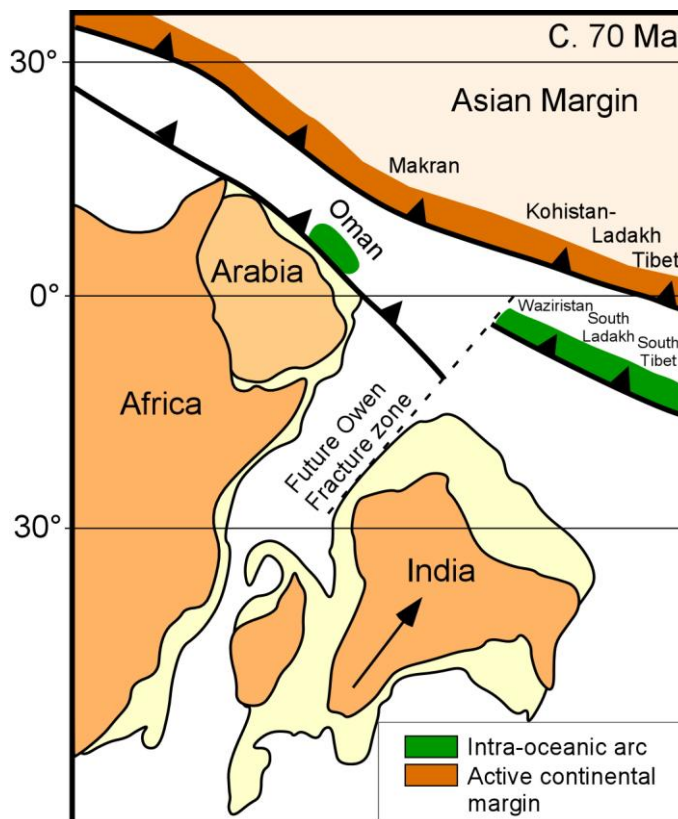


Figure 3: Possible relationships between the Indian plate and the north Tethys subduction during the Upper Cretaceous



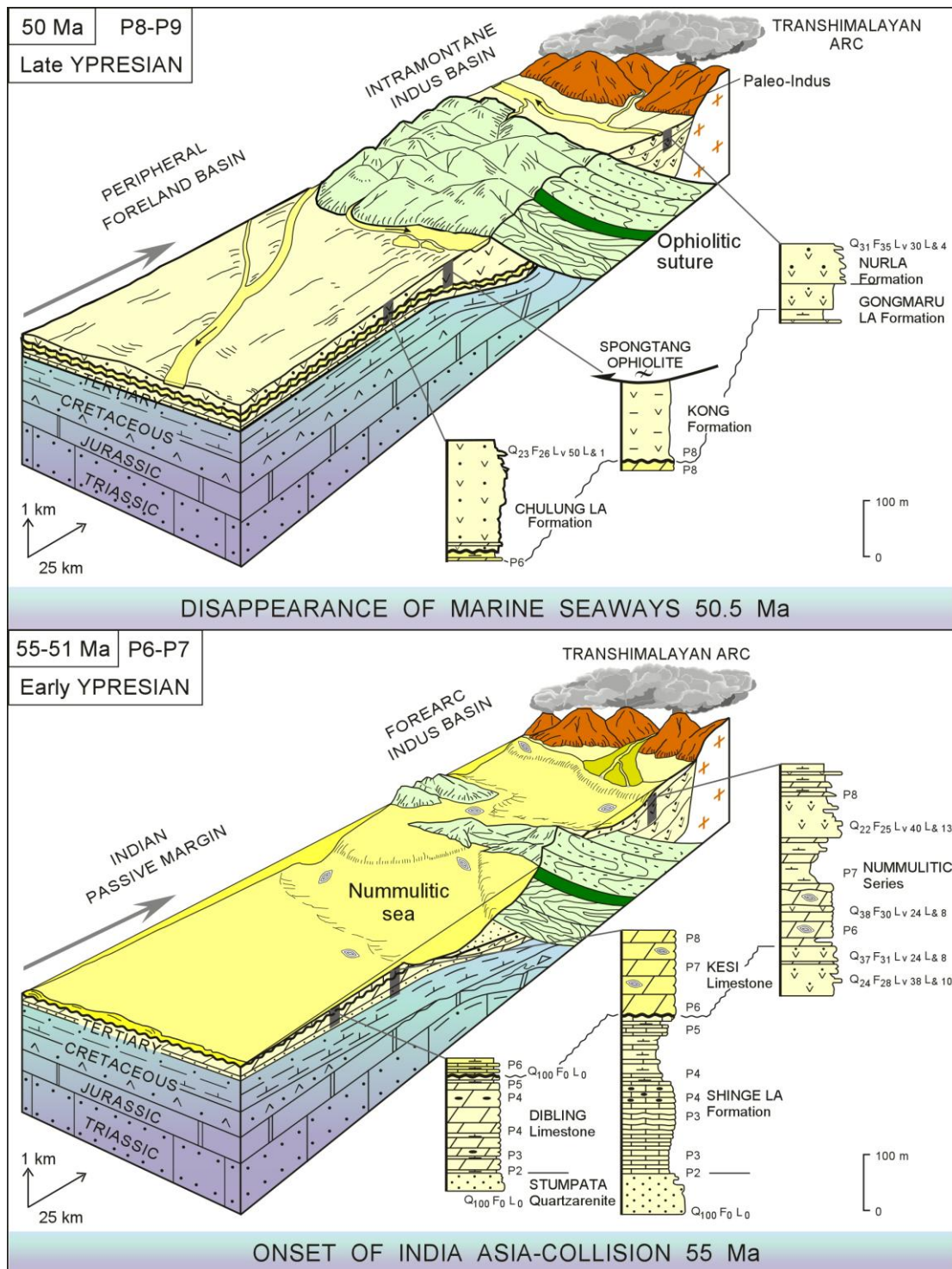


Figure 4: Complete chronology of the India-Asia collision is faithfully recorded in the Paleogene sedimentary record of both Indian passive and Transhimalayan active margins (Garzanti et al., 1987; Garzanti and Van Haver, 1988). After initial continent-continent contact close to the Paleocene-Eocene boundary (55 Ma), the space between colliding continents was filled by shallow Nummulitic seas between 55 and 51 Ma (foraminiferal zones P6-P7; Berggren et al., 1995). At 50.5 Ma (zone P8) marine sedimentation was finally replaced by widespread deposition of continental volcanoclastic redbeds including ophiolitic detritus, which overlapped both the passive-margin succession on the Indian side and the forearc-basin succession on the Asian side. QFL data according to the Gazzi-Dickinson method (Q= quartz; F= feldspars; Lv= volcanic grains; L= other lithic grains; Garzanti et al., 1996).

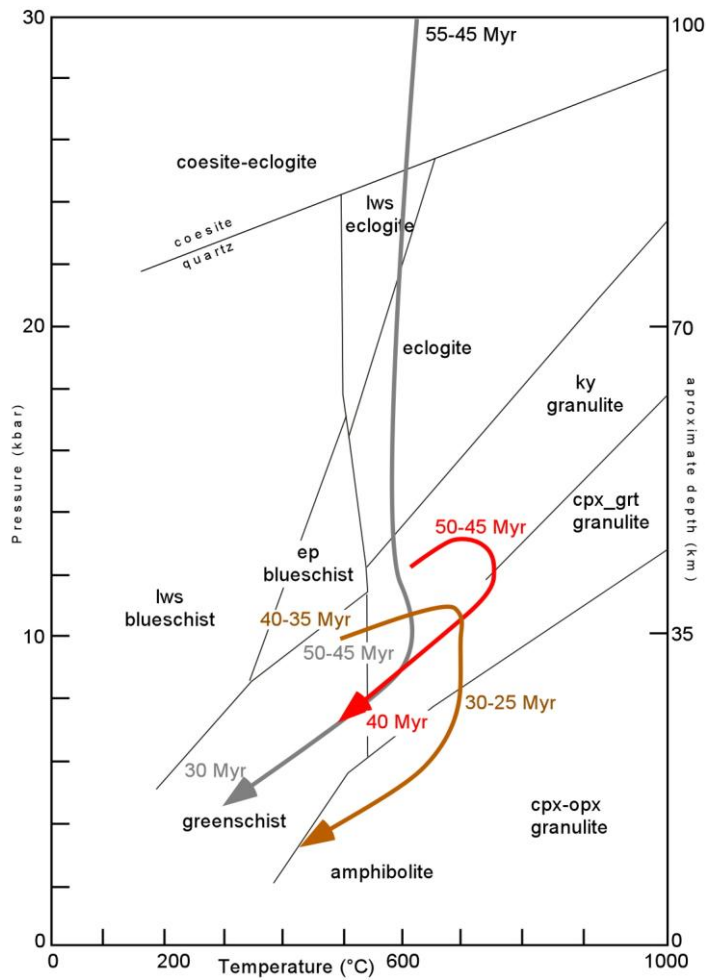


Figure 5: Pressure-Temperature-time (P-T-t) path evolution of the main tectono-metamorphic units in NW Himalaya and South Karakorum. In gray: eclogitic units. In red: the early granulitic unit preserved in NW Pakistan. In brown: the HHC. In orange: the Karakorum metamorphic complex (references in the text).

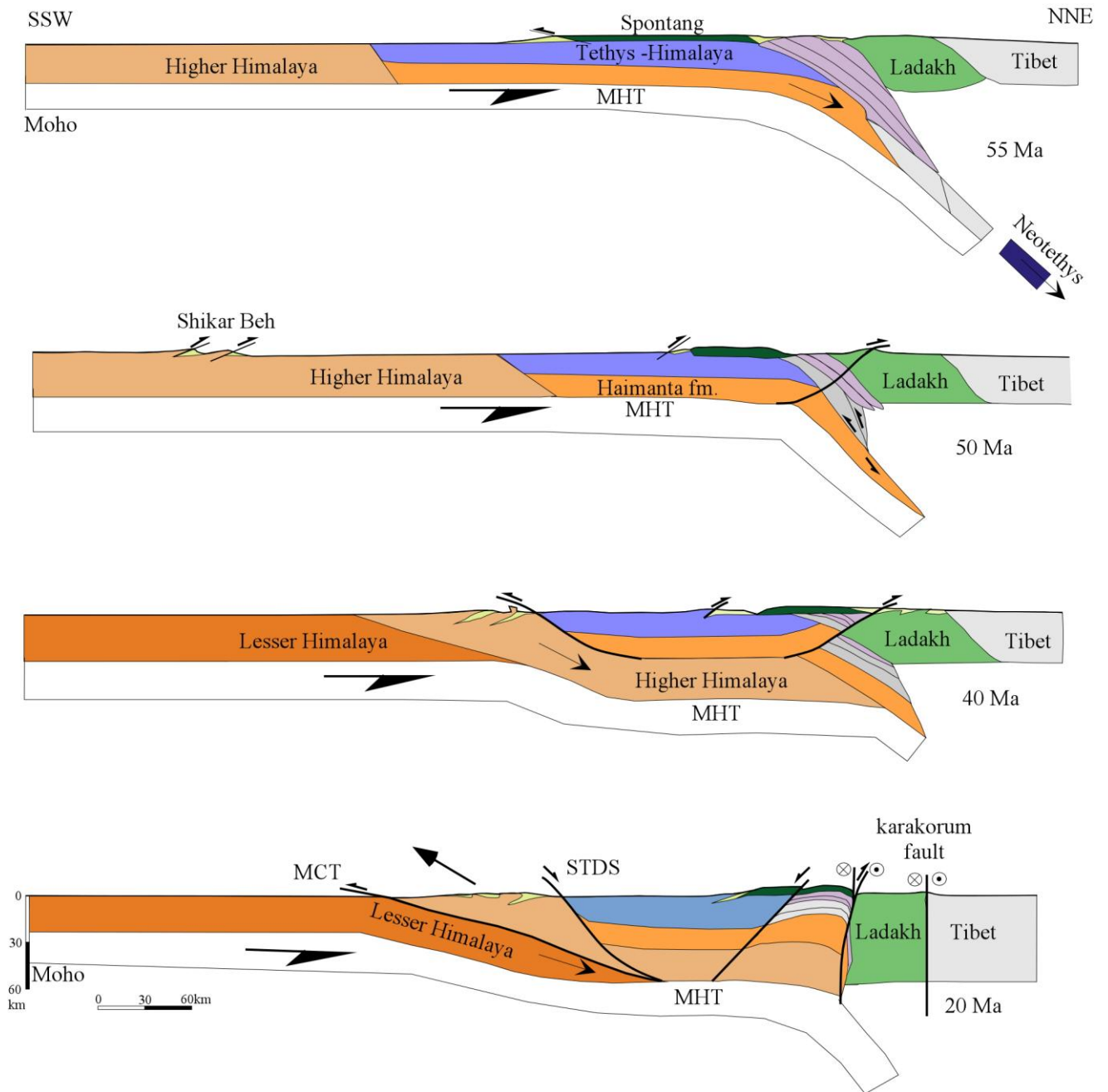


Figure 6 : Balanced cross-sections of the evolution of the Himalayan belt, obtained by retro-deformation of the figure 2. In this reconstruction, we admitted that the MHT is active since the onset of the collision. This model does not take into account the internal deformation, evaluated at about 30% in the ductile crust [Grujic et al., 2000].

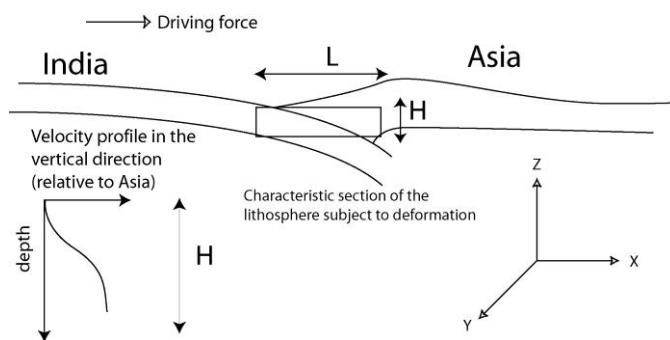


Figure 7: Schematic cross-section of the geometry of the India-Asia collision considered as a zone of two converging block system

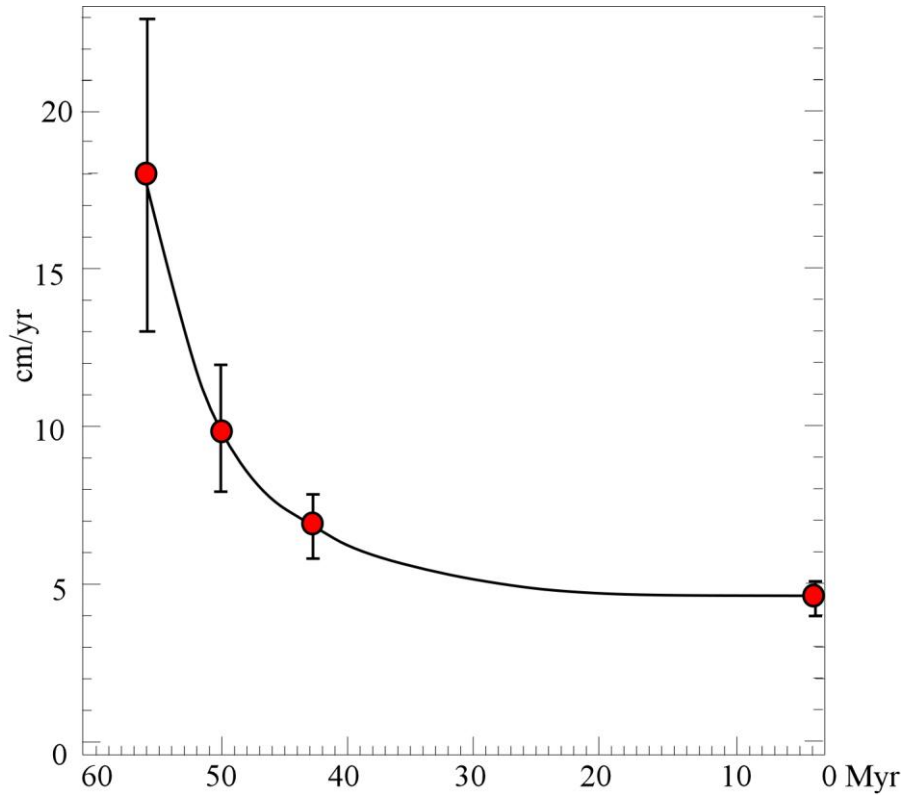


Figure 8: Fitting curve of the north-south velocity of the India-Asia convergence since 55 Ma deduced from the paleomagnetic data (see text for the origin of the data and for the method of fitting)

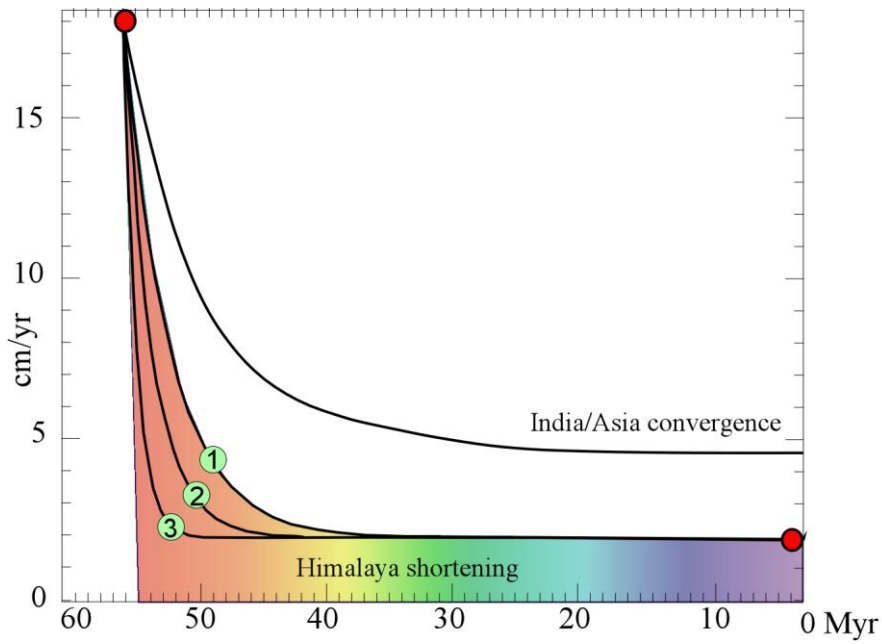


Figure 9: Fitting curves of the Himalayan shortening rate. Curve 1: the Himalayan shortening is fixed to 1600 km. Curve 2: the Himalayan shortening is fixed to 1350 km. Curve 3: the Himalayan shortening is fixed to 1100 km.



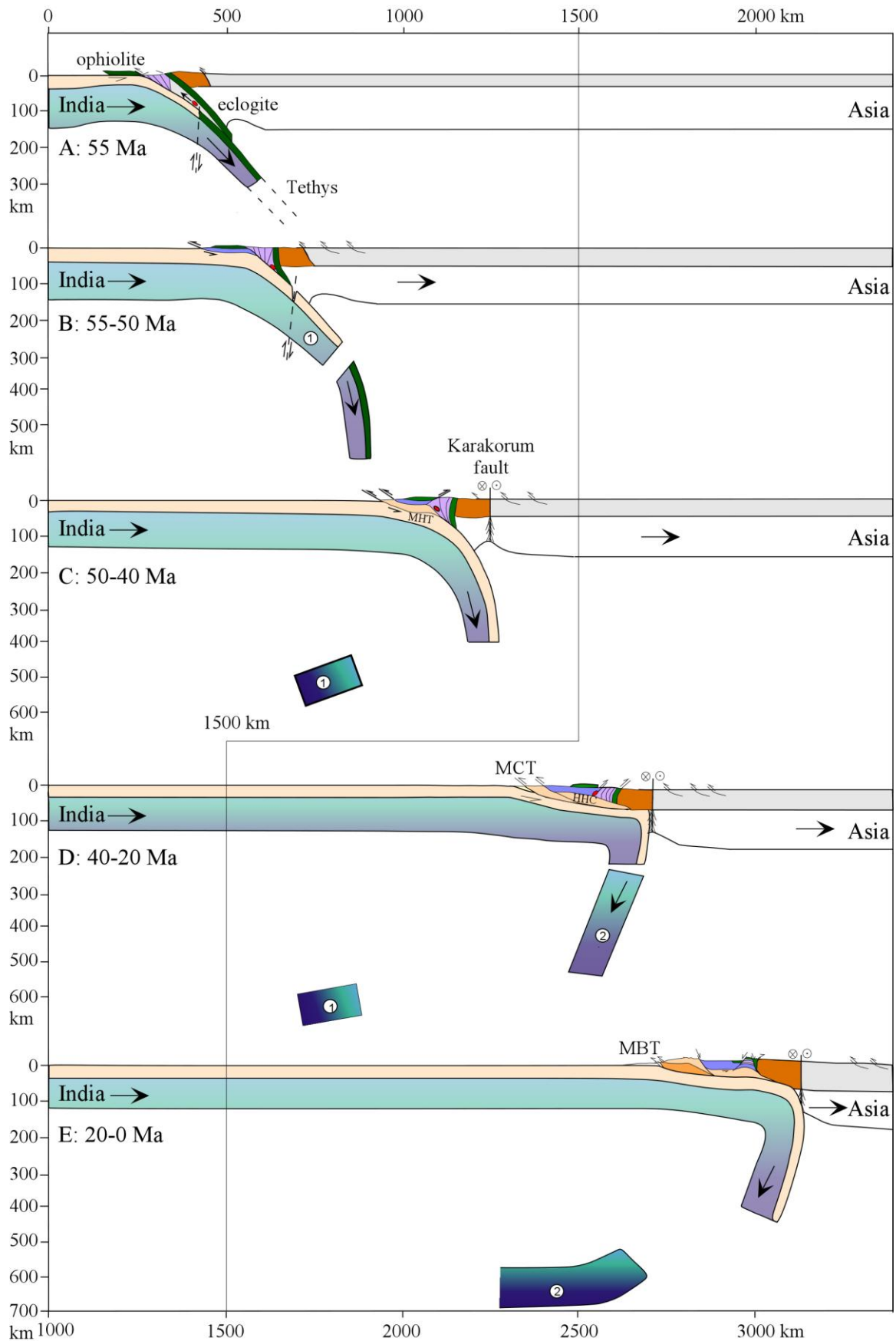


Figure 10: Possible evolution of the NW Himalayan belt at the lithospheric scale. This model takes into account the Tomographic data of Van der Voo et al., [1999], the tectonic evolution and sedimentary discussed in the text and the estimates of Himalayan shortening presented above.

	a	b	c	d
Mean	3.76 cm.yr <sup>-1</sup>	4.7 Myr	0.180 Myr <sup>-1</sup>	4.5 cm.yr <sup>-1</sup>
Fast	0.99 cm.yr <sup>-1</sup>	3.8 Myr	0.178 Myr <sup>-1</sup>	5.0 cm.yr <sup>-1</sup>
Slow	5.58 cm.yr <sup>-1</sup>	5.2 Myr	0.184 Myr <sup>-1</sup>	4.0 cm.yr <sup>-1</sup>

Table 1: Computed values of the parameters used to determine the India-Asia convergence and the associated velocity

Period	Total convergence	Rates
55 Ma -> 50 Ma	670 ± 166 km	13.4 ± 3.3 cm.yr <sup>-1</sup>
50 Ma -> 40 Ma	703 ± 116 km	7.0 ± 1.2 cm.yr <sup>-1</sup>
40 Ma -> 20 Ma	945 ± 110 km	4.7 ± 0.6 cm.yr <sup>-1</sup>
20 Ma -> 0 Ma	897 ± 103 km	4.5 ± 0.5 cm.yr <sup>-1</sup>
55 Ma -> 0 Ma	3215 ± 496 km	5.8 ± 0.9 cm.yr <sup>-1</sup>
45 Ma -> 0 Ma	2140 ± 271 km	4.7 ± 0.6 cm.yr <sup>-1</sup>

Table 2: North-south India/Asia convergence and associated rates computed with the data of the Table 1

	a	b	c	d
Mean	0.82 cm.yr <sup>-1</sup>	5.0 Myr	0.6 Myr <sup>-1</sup>	2.0 cm.yr <sup>-1</sup>

Table 3: Estimated parameters used to calculate Himalayan shortening

Period	Convergence	Velocity
55 Ma -> 50 Ma	345 ± 140 km	6.9 ± 2.9 cm.yr <sup>-1</sup>
50 Ma -> 40 Ma	210 ± 70 km	2.1 ± 0.7 cm.yr <sup>-1</sup>
40 Ma -> 20 Ma	400 ± 20 km	2.0 ± 0.1 cm.yr <sup>-1</sup>
20 Ma -> 0 Ma	400 ± 20 km	2.0 ± 0.1 cm.yr <sup>-1</sup>
55 Ma -> 0 Ma	1355 ± 250 km	2.4 ± 0.4 cm.yr <sup>-1</sup>

Table 4: Himalayan shortening and associated velocities per selected periods

Period	Convergence	Velocity
55 Ma -> 50 Ma	325 ± 280 km	6.5 ± 5.6 cm.yr <sup>-1</sup>
50 Ma -> 40 Ma	493 ± 186 km	4.9 ± 1.9 cm.yr <sup>-1</sup>

40 Ma -> 20 Ma	545 ± 115 km	2.7 ± 0.6 cm.yr <sup>-1</sup>
20 Ma -> 0 Ma	497 ± 108 km	2.5 ± 0.5 cm.yr <sup>-1</sup>
55 Ma -> 0 Ma	1860 ± 689 km	3.4 ± 1.25 cm.yr <sup>-1</sup>

Table 5: Asian shortening and associated velocity

Stage	Oceanic Subduction	Continental Subduction	Continental Collision	Steady-State Collision
Age	> 55 Ma	55-50 Ma	50-25 Ma	25-0 Ma
Crust type	oceanic	thin Indian crust	normal Indian crust	normal Indian crust
India-Asia velocity	18 cm.yr-1	18 -> 10 cm.yr-1	10 -> 4.5 cm.yr-1	4.5 cm.yr-1
India/Asia convergence	0 cm.yr-1	670 ± 166 km	1625 ± 220 km	1122 ± 105 km
Himalayan velocity	18 cm.yr-1	18 -> 2.5 cm.yr-1	2.5 -> 2.0 cm.yr-1	2.0 cm.yr-1
Himalayan shortening	0 km	345 km	510 ± 80 km	500 ± 25 km
Asian velocity	0 cm.yr-1	< 5 cm.yr-1	< 5.0 -> 3.0 cm.yr-1	2.6 cm.yr-1
Asian shortening	0 km	< 250 km	493 ± 186	622 ± 110
Subsidence rate	~ 0.02 mm.yr-1	~ 0.1 mm.yr-1	< 0.01 mm.yr-1	> 0.2 mm.yr-1
Sedimentary facies	shelf sedim	deltaic redbeds	uncorformities	fluvial molasse
Sandstone petrography	quartzarenites	volc+ophio detritus	-----	metam detritus
Indian edge metamorphism	none	eclogite facies	amphibolite facies	greenschist facies
HHC metamorphism	none	none	granulite facies	anatexis
Asian Margin	arc volcanism	cont sedim	uplift	uplift

Table 6: From subduction to collision dynamics in Himalaya